

CERTAIN CURVATURE CONDITIONS SATISFIED BY $N(k)$ -QUASI EINSTEIN MANIFOLDS

Prof. Tarini. A. Kotamkar¹, Prof. Dr. Brajendra. Tiwari²

Dept of Mathematics¹

¹Research scholar in Mathematics Dept.

²Head of the department in RKDF university, Bhopal.

Abstract: In present paper $N(k)$ -quasi Einstein manifolds is studied and its existence is Proved by two non-trivial examples. A physical example of an $N(k)$ -quasi-Einstein manifold is given. We study an $N(k)$ -quasi-Einstein manifold satisfying certain curvature conditions like $\tilde{Z}(\xi, X).S = 0$, and $P(\xi, X).C = 0$. Also studied Ricci-pseudosymmetric $N(k)$ -quasi-Einstein manifolds.

Keywords: Quasi Einstein manifold, $N(k)$ -quasi Einstein manifold, projective curvature tensor, concircular curvature tensor, conformal curvature tensor, Ricci-pseudosymmetric manifold.

I. INTRODUCTION

An Einstein manifold is a Riemannian or a semi-Riemannian manifold (M_n, g) , $n = \dim M \geq 2$, if the following condition holds on M , where S is the Ricci tensor and r denotes the scalar curvature of (M_n, g)

$$(1.1) \quad S = \frac{r}{n} g$$

(1.1) is called the Einstein metric condition as mentioned in [4]. In Riemannian geometry as well as in general theory of relativity Einstein manifolds play an important role. By a curvature condition imposed on their Ricci tensor [4], Einstein manifolds form a natural sub-class of various classes of Riemannian or semi-Riemannian manifolds. The class of Riemannian manifolds (M_n, g) realizing the following relation consists of every Einstein manifold.

$$(1.2) \quad S(X, Y) = ag(X, Y) + b\eta(X)\eta(Y)$$

where a, b are smooth functions and η is a non-zero 1-form such that

$$(1.3) \quad g(X, \xi) = \eta(X), \quad g(\xi, \xi) = \eta(\xi) = 1$$

Where X stands for all vector fields.

A quasi Einstein manifold [7] is a non-flat Riemannian manifold (M_n, g) ($n > 2$) if its Ricci tensor S of type (0,2) is not identically zero and the condition in (1.2) is satisfied. The unit vector field ξ is called the generator of the manifold and we call η the associated 1-form.

The study of exact solutions of the Einstein field equations and also during considerations of quasi-umbilical hypersurfaces of semi-Euclidean spaces, Quasi Einstein manifolds arose. In [15], Naschie on the theory of elementary particles and using Einstein's unified field equation, showed the expectation number of elementary particles of the standard model. In [14] Naschie also discussed possible connections between Godel's classical solution of Einstein's field equations and E-infinity. In the general theory of relativity quasi Einstein manifolds have some importance. The Robertson-Walker spacetime are quasi Einstein manifolds by [13]. In general relativity [9], quasi Einstein manifold can be taken as a model of the perfect fluid spacetime.

Chaki [5], Guha [16], De and Ghosh [10, 11] and many others have studied quasi Einstein manifolds. In [20] Ozgur studied

super quasi-Einstein manifolds and generalized quasi-Einstein manifolds [18]. In [17] Nagaraja studied $N(k)$ -mixed quasi-Einstein manifolds.

The Riemannian curvature tensor of a Riemannian manifold M be denoted by R . In [23] the k -nullity distribution $N(k)$ of a Riemannian manifold M is defined by

$$N(k) \cdot p \rightarrow N_p(k) = \{ Z \in T_p M : R(X, Y)Z = k[g(Y, Z)X - g(X, Z)Y] \},$$

Where k is some smooth function. If the generator ξ belongs to some k -nullity distribution $N(k)$, in a quasi Einstein manifold M , then M is said to be a $N(k)$ -quasi Einstein manifold by [24]. By the following k is not arbitrary:

Lemma 2.1. [22] it follows that in an n -dimensional $N(k)$ -quasi Einstein manifold

$$(1.4) \quad k = [(a + b)/(n-1)]$$

Now, it is immediate to note that in an n -dimensional $N(k)$ -quasi-Einstein manifold [22]

$$(1.5) \quad R(X, Y)\xi = [(a + b)/(n-1)] [\eta(Y)X - \eta(X)Y],$$

which is equivalent to

$$R(X, \xi)Y = [(a + b)/(n-1)] [\eta(Y)X - g(X, Y)\xi] = -R(\xi, X)Y.$$

From (1.5) we get

$$R(\xi, X)\xi = [(a + b)/(n-1)] [\eta(X)\xi - X].$$

In [24], it was shown that an n -dimensional conformally flat quasi Einstein manifold is an $N[(a + b)/(n-1)]$ quasi Einstein manifold and in particular a 3-dimensional quasi Einstein manifold is an $N[(a + b)/(2)]$ quasi Einstein manifold. In differential geometry and also in general theory of relativity the conformal curvature tensor plays an important role. Definition of the Weyl conformal curvature tensor C of a Riemannian manifold (M_n, g) ($n > 3$) is :

$$C(X, Y)Z = R(X, Y)Z - [n-1]/2 [g(Y, Z)QX - g(X, Z)QY + S(Y, Z)X - S(X, Z)Y] + [r/(n-1)(n-2)] [g(Y, Z)X - g(X, Z)Y]$$

where Q is the symmetric endomorphism of the tangent space at each point corresponding to the Ricci tensor S that is, $g(QX, Y) = S(X, Y)$ and r is the scalar curvature. For the dimension $n = 3$, then the conformal curvature tensor vanishes identically. The projective curvature tensor P and the concircular curvature tensor \check{Z} in a Riemannian manifold (M_n, g) are defined by [25]

$$P(X, Y)W = R(X, Y)W - [1/(n-1)] [S(Y, W)X - S(X, W)Y],$$

$$\check{Z}(X, Y)W = R(X, Y)W - [r/(n-1)] [g(Y, W)X - g(X, W)Y], \text{ respectively.}$$

It is proved in [24] by the authors that conformally flat quasi-Einstein manifolds are certain $N(k)$ -quasi-Einstein manifolds. In [23] the derivation conditions $R(\xi, X) \cdot R = 0$ and $R(\xi, X) \cdot S = 0$ have been studied, where R denotes the curvature tensor and S denotes Ricci tensor. Further in [22] "Ozgür and Tripathi studied the derivation conditions $\check{Z}(\xi, X) \cdot R = 0$ and $\check{Z}(\xi, X) \cdot \check{Z} = 0$ on $N(k)$ -quasi-Einstein manifold where \check{Z} is the concircular curvature tensor. It was in [22] proved that $k = [(a+b)/(n-1)]$ for an $N(k)$ -quasi-Einstein manifold. The condition $R \cdot P = 0$, $P \cdot S = 0$ and $P \cdot P = 0$ for an $N(k)$ -quasi-Einstein manifolds, where P denotes the projective curvature tensor, studied by "Ozgür [19] and some physical examples of $N(k)$ -quasi-Einstein manifolds are given. "Ozgür and Sular [21] in 2008, studied $N(k)$ -quasi-Einstein manifold satisfying $R \cdot C = 0$ and $R \cdot \check{C} = 0$, where C represents the conformal curvature tensor and \check{C} represents the quasi-conformal

curvature tensor. In continuation of previous studies this paper is presented. In this paper after preliminaries some examples of $N(k)$ -quasi-Einstein manifolds are discussed. In this paper some we give a physical example of an $N(k)$ -quasi-Einstein manifold and we study $N(k)$ -quasi-Einstein manifold satisfying $\check{Z}(\xi, X) \cdot S = 0$ in Section 5. Section 6 deals with $N(k)$ -quasi-Einstein manifolds satisfying $P(\xi, X) \cdot C = 0$. In and Section 7, we study $N(k)$ -quasi-Einstein manifolds satisfying $\check{Z}(\xi, X) \cdot C = 0$. Finally, we study Ricci-pseudosymmetric $N(k)$ -quasi-Einstein manifolds.

II. PRELIMINARIES

From (1.2) and (1.3) it follows that

$$r = an + b \text{ and } QX = (a + b)X,$$

$$S(X, \xi) = k(n - 1) \eta(X),$$

Where r is the scalar curvature and Q is the Ricci operator.

In an n -dimensional $N(k)$ -quasi-Einstein manifold M , the projective curvature tensor P , the concircular curvature tensor \check{Z} and the conformal curvature tensor C satisfy the following relations:

$$P(X, Y) \xi = 0,$$

$$(2.1) \quad \eta(P(\xi, X)Y) = \frac{b}{[n-1]} [g(X, Y) \xi - \eta(X) \eta(Y) \xi],$$

$$(2.2) \quad P(X, Y)Z = \frac{b}{[n-1]} [g(Y, Z) \eta(X) - g(X, Z) \eta(Y)],$$

$$(2.3) \quad Z(X, Y)Z = [k - \frac{r}{n(n-1)}][g(Y, Z)X - g(X, Z)Y],$$

$$(2.4) \quad Z(\xi, Y)Z = [k - \frac{r}{n(n-1)}][g(Y, Z) \xi - \eta(Z)Y],$$

$$(2.5) \quad C(X, Y)Z = \frac{k - \frac{r}{(n-1)(n-2)}}{[1/(n-2)]} [g(Y, Z)X - g(X, Z)Y] - [(2a + b)g(Y, Z)X - (2a + b)g(X, Z)Y + b \eta(Y) \eta(Z)X - b \eta(X) \eta(Z)Y],$$

$$(2.6) \quad \eta(C(X, Y)Z) = 0,$$

$$(2.7) \quad C(\xi, Y)Z = -\frac{b}{n-2} [\eta(Y) \eta(Z) \xi - \eta(Z)Y],$$

for all vector fields X, Y, Z on M .

III. EXAMPLES OF $N(K)$ -QUASI EINSTEIN MANIFOLDS

Example 3.1. An $N(k)$ -quasi Einstein manifold consists of special para-Sasakian manifold with vanishing D-concircular curvature tensor V . A Riemannian manifold, M_n having a unit concircular vector field ξ such that $\nabla_X \xi = \varepsilon(-X + \eta(X)\xi)$, $\eta(X) = g(X, \xi)$, $\varepsilon = \pm 1$, then M_n is called a special para-Sasakian manifold [1–3]. The notion of a D-concircular curvature tensor V is recently introduced by Chuman [8]. V is given by the following equation

$$V(X, Y, Z, W) = \check{R}(X, Y, Z, W) + \frac{r + 2(n-1)}{(n-1)(n-2)} [g(X, Z)g(Y, W) - g(Y, Z)g(X, W)]$$

$$(3.1) \quad -\frac{r+n(n-1)}{(n-1)(n-2)} [g(X, Z) \eta(Y) \eta(W) - g(Y, Z) \eta(X) \eta(W) + g(Y, W) \eta(Z) \eta(X) - g(X, W) \eta(Y) \eta(Z)],$$

Where $\check{R}(X, Y, Z, W) = g(R(X, Y)Z, W)$ for $R(X, Y)Z$ the curvature tensor of type (1, 3). If $V(X, Y, Z, W) = 0$, then from (3.1) it follows that

$$(3.2) \quad \check{R}(X, Y, Z, W) = \frac{r+2(n-1)}{(n-1)(n-2)} [g(Y, Z) g(X, W) - g(X, Z) g(Y, W)] + \frac{r+n(n-1)}{(n-1)(n-2)} [g(X, Z) \eta(Y) \eta(W) \eta(g(Y, Z) \eta(X) \eta(W) + g(Y, W) \eta(Z) \eta(X) - g(X, W) \eta(Y) \eta(Z)],$$

Putting $X = W = e_i$ in (3.2) where $\{e_i\}$ is an orthonormal basis of the tangent space at each point of the manifold and taking summation over $i, 1 \leq i \leq n$, we get

$$S(Y, Z) = a g(Y, Z) + b \eta(Y) \eta(Z),$$

where $a = \frac{r+n-1}{n-1}$ and $b = -\frac{r+n(n-1)}{n-1}$. Therefore, $\frac{a+b}{n-1} = -1$. Hence a special para-Sasakian manifold with vanishing D-concircular curvature tensor is an $N(-1)$ -quasi-Einstein manifold.

IV. A PHYSICAL EXAMPLE OF AN $N(K)$ -QUASI-EINSTEIN MANIFOLD

Example 4.1 An $N(k)$ -quasi-Einstein manifold in general relativity by the coordinate free method of differential geometry is concerned with this example. In this method of study the space time of general relativity is regarded as a connected four-dimensional semi-Riemannian manifold (M_4, g) with Lorentzian metric g with signature $(-, +, +, +)$. With the study of causal character of vectors of the manifold begins the geometry of the Lorentzian manifold. The Lorentzian manifold becomes a convenient choice for the study of general relativity due to this causality.

Here we consider a perfect fluid $(PRS)_4$ spacetime of non-zero scalar curvature and having the basic vector field U as the time like vector field of the fluid, that is, $g(U, U) = -1$. An n -dimensional semi-Riemannian manifold is said to be pseudo Ricci-symmetric [6] if the Ricci tensor S satisfies the condition

$$(\nabla_X S)(Y, Z) = 2A(X)S(Y, Z) + A(Y)S(X, Z) + A(Z)S(Y, X).$$

Such a manifold is denoted by $(PRS)_n$.

For the perfect fluid spacetime, we have the Einstein equation without cosmological constant as

$$(4.1) \quad S(X, Y) - \frac{1}{2} r g(X, Y) = \kappa T(X, Y),$$

where κ is the gravitational constant, T is the energy-momentum tensor of type (0, 2) given by

$$(4.2) \quad T(X, Y) = (\sigma + p)B(X)B(Y) + pg(X, Y),$$

with σ and p as the energy density and isotropic pressure of the fluid respectively.

Using (4.2) in (4.1) we get

$$(4.3) \quad S(X, Y) - 12rg(X, Y) = \kappa[(\sigma + p)B(X)B(Y) + pg(X, Y)].$$

Taking a frame field and contracting (4.3) over X and Y we have

$$(4.4) \quad r = \kappa(\sigma - 3p).$$

Using (4.4) in (4.3) yields

$$(4.5) \quad S(X, Y) = \kappa[(\sigma + p)B(X)B(Y) + \frac{(\sigma - p)}{2}g(X, Y)].$$

Putting $Y = U$ in (4.5) and since $g(U, U) = -1$, we get

$$(4.6) \quad S(X, U) = -\frac{\kappa}{2}[\sigma + 3p]B(X).$$

Again for $(PRS)_n$ spacetime [6], $S(X, U) = 0$. This condition will be satisfied by the equation (4.6) if

$$(4.7) \quad \sigma + 3p = 0 \text{ as } \kappa \neq 0 \text{ and } A(X) \neq 0.$$

Using (4.4) and (4.7) in (4.5) we see that

$$S(X, Y) = \frac{r}{3}[B(X)B(Y) + g(X, Y)].$$

Thus we can state the following:

An $N(\frac{2r}{9})$ -quasi-Einstein manifold is a perfect fluid pseudo Ricci-symmetric spacetime.

V. $N(K)$ -QUASI EINSTEIN MANIFOLD SATISFYING $\check{Z}(\xi, X) \cdot S = 0$

In this section we consider an n -dimensional $N(k)$ -quasi-Einstein manifold M satisfying the condition

$$(\check{Z}(\xi, X) \cdot S)(Y, Z) = 0.$$

Putting $Z = \xi$ we get

$$(5.1) \quad S(\check{Z}(\xi, X)Y, \xi) + S(Y, \check{Z}(\xi, X)\xi) = 0.$$

Using (1.2), (2.3) and (2.4) in (5.1) we get

$$(5.2) \quad S(\check{Z}(\xi, X)Y, \xi) = \left[k - \frac{r}{n(n-1)}\right](a+b)[g(X, Y) - \eta(X)\eta(Y)],$$

and

$$(5.3) \quad S(Y, \check{Z}(\xi, X)\xi) = \left[k - \frac{r}{n(n-1)}\right][(a+b)\eta(X)\eta(Y) - S(X, Y)].$$

Using (5.2) and (5.3) in (5.1), we obtain

$$\left[k - \frac{r}{n(n-1)}\right][(a+b)g(X, Y) - S(X, Y)] = 0.$$

Therefore, either the scalar curvature of M is $kn(n-1)$ or, $S = (a+b)g$ which implies that M is an Einstein manifold. But this contradicts the definition of quasi-Einstein manifold. The converse is trivial.

Thus we can state the following:

Theorem 5.1. The condition $\check{Z}(\xi, X) \cdot S = 0$ is satisfied by an n-dimensional N(k)-quasi-Einstein manifold M if and only if the scalar curvature is $kn(n-1)$.

VI. N(K)-QUASI-EINSTEIN MANIFOLD SATISFYING $P(\xi, X) \cdot C = 0$

In this section we consider an n-dimensional N(k)-quasi-Einstein manifold satisfying the condition

$$(P(\xi, X) \cdot C)(Y, Z)W = 0$$

Then we have

$$(6.1) \quad P(\xi, X)C(Y, Z)W - C(P(\xi, X)Y, Z)W - C(Y, P(\xi, X)Z)W - C(Y, Z)P(\xi, X)W = 0.$$

Using (2.1) in (8.1) we obtain

$$\frac{b}{n-1} [g(X, C(Y, Z)W) \xi - \eta(X) \eta(C(Y, Z)W) \xi - C(g(X, Y) \xi - \eta(X) \eta(Y) \xi, Z)W - C(Y, g(X, Y) \xi - \eta(X) \eta(Z) \xi)W - C(Y, Z)(g(X, W) \xi - \eta(X) \eta(W) \xi)] = 0,$$

which implies either $b = 0$, or

$$(6.2) \quad g(X, C(Y, Z)W) \xi - \eta(X) \eta(C(Y, Z)W) \xi - g(X, Y)C(\xi, Z)W + \eta(X) \eta(Y)C(\xi, Z)W - g(X, Z)C(Y, \xi)W + \eta(X) \eta(Z)C(Y, \xi)W - g(X, W)C(Y, Z) \xi + \eta(X) \eta(W)C(Y, Z) \xi = 0,$$

holds on M. Since $b \neq 0$, hence (8.2) holds.

Taking inner product with ξ and using (2.5)–(2.7) in (8.2) we obtain $g(X, C(Y, Z)W) = 0$.

Thus we can state the following:

Theorem 6.1. An n-dimensional N(k)-quasi-Einstein manifold M satisfies the condition $P(\xi, X) \cdot C = 0$ if and only if the manifold is conformally flat.

VII. N(K)-QUASI-EINSTEIN MANIFOLD SATISFYING $\check{Z}(\xi, X) \cdot C = 0$

In this section we consider an n-dimensional N(k)-quasi-Einstein manifold satisfying the condition

$$(\check{Z}(\xi, X) \cdot C)(Y, Z)W = 0.$$

Then we have

$$(7.1) \quad \check{Z}(\xi, X)C(Y, Z)W - C(\check{Z}(\xi, X)Y, Z)W - C(Y, \check{Z}(\xi, X)Z)W - C(Y, Z)(\check{Z}(\xi, X)W) = 0.$$

Using (2.4) in (9.1) we obtain μ

$$(7.2) \quad k - r n(n-1) [g(X, C(Y, Z)W) \xi - \eta(C(Y, Z)W)X - g(X, Y)C(\xi, Z)W \\ + \eta(Y)C(X, Z)W - g(X, Z)C(Y, \xi)W + \eta(Z)C(Y, X)W \\ - g(X, W)C(Y, Z)\xi + \eta(W)C(Y, Z)X] = 0,$$

which gives either $r = kn(n-1)$ or

$$g(X, C(Y, Z)W) \xi - \eta(C(Y, Z)W)X - g(X, Y)C(\xi, Z)W + \eta(Y)C(X, Z)W \\ - g(X, Z)C(Y, \xi)W + \eta(Z)C(Y, X)W - g(X, W)C(Y, Z)\xi + \eta(W)C(Y, Z)X = 0.$$

Taking inner product with ξ and using (2.5)–(2.7) in (9.2) we obtain

$$g(X, C(Y, Z)W) = 0.$$

Thus we can state the following.

Theorem 7.1. The condition $\tilde{Z}(\xi, X) \cdot C = 0$ is satisfied by an n -dimensional $N(k)$ -quasi-Einstein manifold M if and only if the manifold is conformably flat.

VIII. RICCI-PSEUDOSYMMETRIC $N(K)$ -QUASI-EINSTEIN MANIFOLDS

An n -dimensional Riemannian manifold (M_n, g) is called Ricci-pseudosymmetric[12] if the tensors $R \cdot S$ and $Q(g, S)$ are linearly dependent, where

$$(8.1) \quad (R(X, Y) \cdot S)(Z, W) = -S(R(X, Y)Z, W) - S(Z, R(X, Y)W),$$

$$(8.2) \quad Q(g, S)(Z, W, X, Y) = -S((X \wedge Y)Z, W) - S(Z, (X \wedge Y)W),$$

and

$$(8.3) \quad (X \wedge Y)Z = g(Y, Z)X - g(X, Z)Y, \quad (10.3)$$

for vector fields X, Y, Z, W on M_n .

The condition of Ricci-pseudosymmetry is equivalent to

$$(R(X, Y) \cdot S)(Z, W) = L_s Q(g, S)(Z, W, X, Y),$$

holding on the set

$$U_s = \{x \in M : S \neq \frac{r}{n} \text{ at } x\},$$

where L_s is some function on U_s . If $R \cdot S = 0$ then M_n is called Ricci-semisymmetric. Every Ricci-semisymmetric manifold is Ricci-pseudosymmetric but the converse is not true [12].

Let us assume that the manifold under consideration is Ricci-pseudosymmetric. Then with the help of (10.1)–(10.3) we can write

$$(8.4) \quad S(R(X, Y)Z, W) + S(Z, R(X, Y)W) = L_s \{g(Y, Z)S(X, W) - g(X, Z)S(Y, W) \\ + g(Y, W)S(X, Z) - g(X, W)S(Y, Z)\}.$$

Using (1.2) and (1.4) in (10.4) we obtain

$$(8.5) \quad \left[\frac{b(a+b)}{n-1} - bL_s \right] \{g(Y,Z)\eta(X)\eta(W) - g(X,W)\eta(Y)\eta(Z) + g(Y,W)\eta(X)\eta(Z) - g(X,Z)\eta(Y)\eta(W)\} = 0.$$

Putting $Y = Z = \xi$ in (10.5) we have

$$(8.6) \quad \left[\frac{b(a+b)}{n-1} - bL_s \right] \eta(X)\eta(W) - g(X,W) = 0.$$

Again putting $X = W = e_i$ in (10.6), where $\{e_i\}$, $(i = 1, 2, \dots, n)$ is an orthonormal basis of the tangent space at any point of the manifold and then taking the sum for $1 \leq i \leq n$, we obtain

$$b(a+b)n - 1 - bL_s(1-n) = 0,$$

which implies that $L_s = \frac{a+b}{n-1}$. Thus we can state the following:

Theorem 8.1. If $a + b = 0$ then a Ricci-pseudosymmetric $N(k)$ -quasi-Einstein manifold is a Ricci-semisymmetric manifold.
condition is $\check{Z}(\xi, X) \cdot S = 0$ if and only if the scalar curvature is $kn(n-1)$. the condition $\check{Z}(\xi, X) \cdot S = 0$ if and only if the scalar

IX. CONCLUSION

During the study of exact solutions of the Einstein field equations as well as during considerations of quasi-umbilical hyper surfaces of semi-Euclidean spaces arose Quasi-Einstein manifolds. In the introduction the importance of an $N(k)$ -quasi-Einstein is presented. We prove that a special para Sasakian manifold with vanishing D-concircular curvature tensor V is an $N(k)$ -quasi-Einstein manifold. A physical example of $N(k)$ -quasi-Einstein manifolds is given. $N(k)$ -quasi-Einstein manifolds satisfying the curvature conditions has been considered $\check{Z}(\xi, X) \cdot S = 0$, $P(\xi, X) \cdot \check{Z} = 0$, $\check{Z}(\xi, X) \cdot P = 0$, $P(\xi, X) \cdot C = 0$ and $\check{Z}(\xi, X) \cdot C = 0$, where P , \check{Z} , C , S are projective curvature tensor, concircular curvature tensor, conformal curvature tensor and Ricci tensor respectively. Finally we prove that a Ricci-pseudosymmetric $N(k)$ -quasi-Einstein manifold is a Ricci-semi symmetric manifold under certain condition.

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